



VSSEC

Victorian Space Science Education Centre

ビクトリア宇宙教育センター

ビクトリア州の学生と教員をSTEM教育で支援



火星 ミッション

- 中学生対象
- シナリオ型のアクティブラーニング



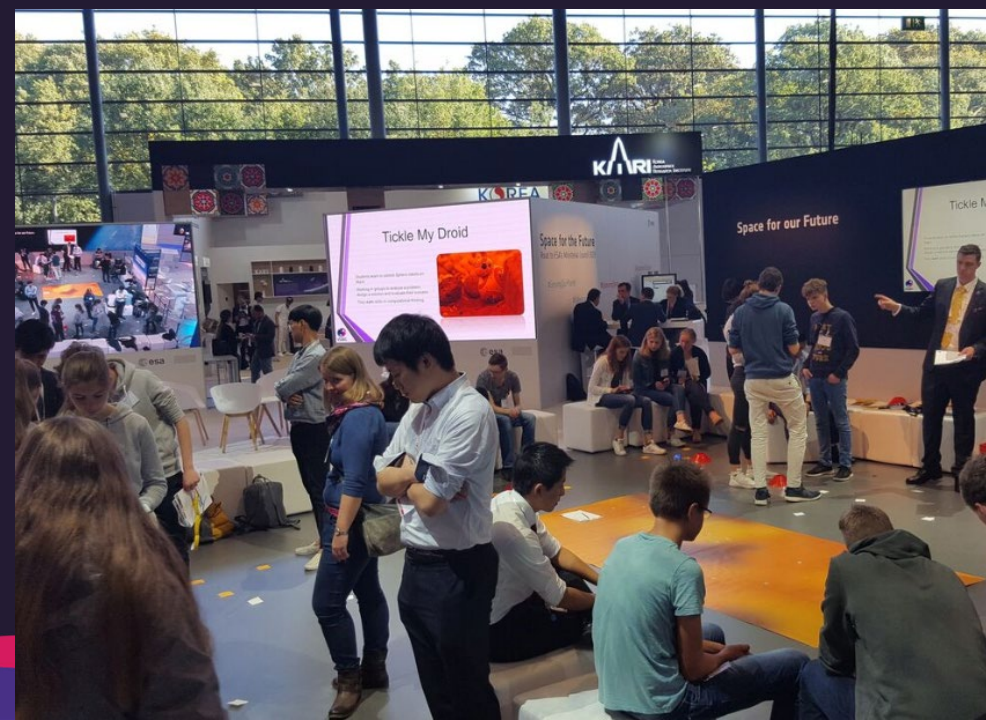
教員研修

- 2000の従来からのプログラム
- 訪問校のすべての教師が対象
- 教育方法を扱う
- 「科学教育のための教育の科学」



国際パートナーシップ

- 国際宇宙教育会議(ISEB)の準会員
- VSSEC は国際宇宙教育コミュニティにおいてオーストラリアを代表
- 国際宇宙会議 (IAC) に参加者及び登壇者として参加



電波天文学：目に見えない宇宙



View

Mount Control

Mount type: WP-100

COM Port: COM3

Actual Antenna Coordinates

30h 22m 29.7s L. Sidereal Time
 10h 53m 19.0s Universal Time
 08h 45m 59.6s Right Ascension
 -52° 53' 34.5" Declination
 +133° 55' 48.0" Azimuth
 +60° 36' 00.0" Elevation
 -88° 57' 38.0" Galactic Long
 06° 02' 42.3" Galactic Lat

Link:
 AZ:
 EL:
 Park:
 Track:
 Slew:
 Coord. Syst: Equ Gal

GOTO: OnA

Equ AltAz Gal

hour min sec
 08 41 09.7 Right Ascension
 deg min sec
 -53 06 35 Declination

GOTO

Commanded Antenna Coordinates

08h 41m 09.7s Right Ascension
 52° 53' 25.0" Declination
 00h 03m 22.6s Error on RA
 -00° 00' 09.5" Error on Dec
 +134° 13' 44.9" Azimuth
 +61° 02' 53.8" Elevation
 -00° 17' 56.9" Error on Az
 -00° 26' 53.8" Error on D

STOP

TRACK

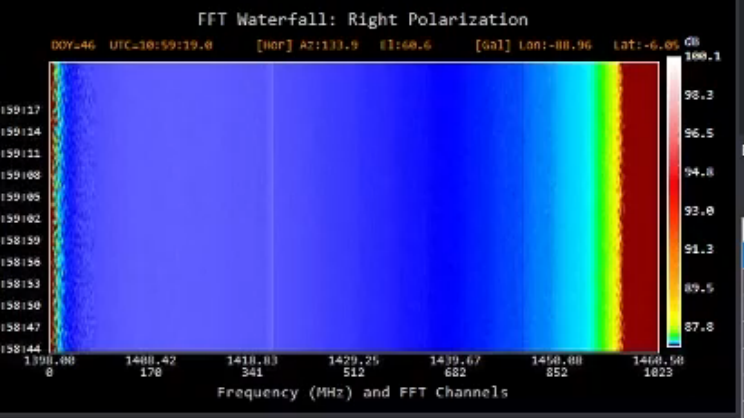
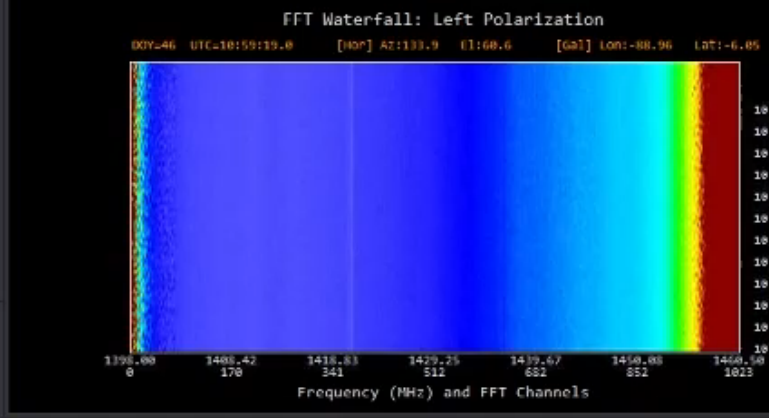
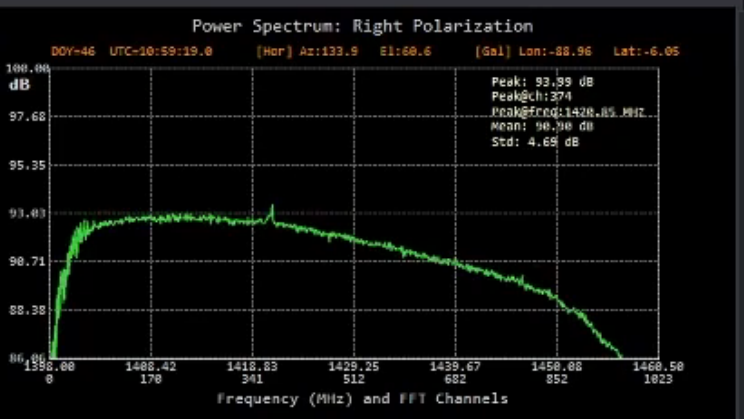
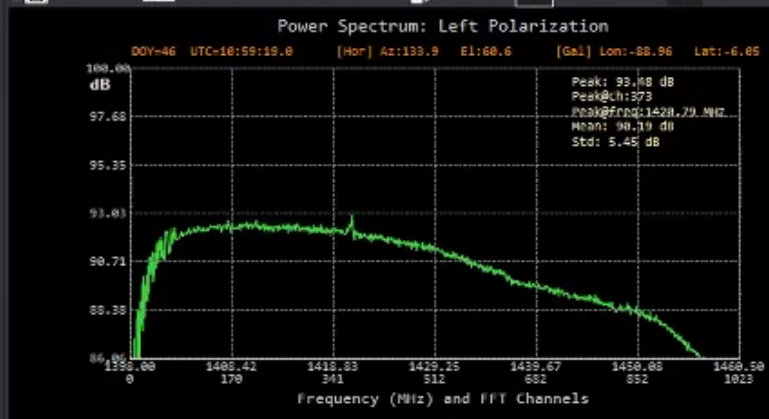
PARK

Unwrapping Monitor

Az: 133.93°
 HA: 21.61°

Status: Cable wrap OK

Rates
 Az: -72.00%/s E: -0.29%/s



```
2021.046.10:59:13.36>/approximating To target...
2021.046.10:59:13.34>/antenna slewing...
2021.046.10:59:13.34>/approximating To target...
2021.046.10:59:14.34>/antenna slewing...
2021.046.10:59:14.35>/approximating To target...
2021.046.10:59:15.35>/antenna slewing...
2021.046.10:59:15.35>/approximating To target...
2021.046.10:59:16.34>/antenna slewing...
2021.046.10:59:16.34>/approximating To target...
2021.046.10:59:17.35>/antenna slewing...
2021.046.10:59:17.35>/approximating To target...
2021.046.10:59:18.35>/antenna slewing...
2021.046.10:59:18.35>/approximating To target...
2021.046.10:59:19.34>/antenna slewing...
2021.046.10:59:19.34>/approximating To target...
```

(Please type here a command line. Type 'help' for list of commands)

Receivers and Backend Control

Backend Status: LINK RX ERR, LOCAL CALON

H142-One

IF and BBC Total Power Levels

[USB] Lower Side Band

[USB] Upper Side Band

IF and Sampler Settings

Center of Band (Sky Freq): 1429.25

IF Gain (dB): ia +16, ib +16

Samp.(s): 0.3

LO link:

Baseband Converter Control

ia [LFFT] ia [BBC] ib [RIGHT] ib [BBC]

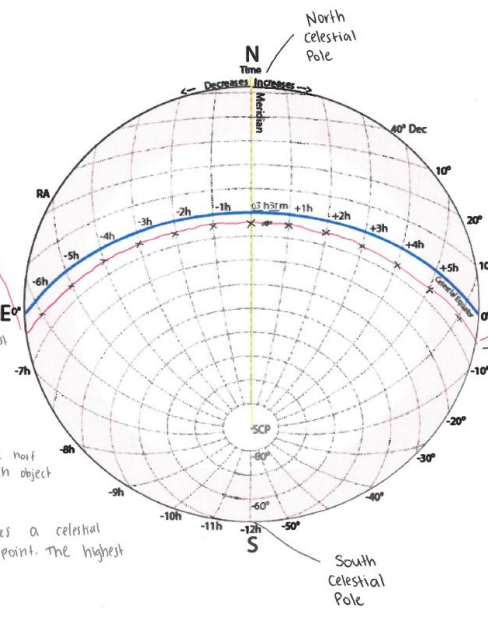
BBC Preset: BBC_setup

L	U	bbc	freq	if	sky	bwu	bwl	over	mode	gain	gain
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1	0	ia	1398.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	2	8	ia	1405.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	3	20	ia	1418.00	4	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	4	24	ia	1422.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	5	32	ia	1430.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	6	40	ia	1438.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	7	48	ia	1446.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	8	56	ia	1454.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	9	0	ib	1398.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	10	8	ib	1405.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	11	20	ib	1418.00	4	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	12	24	ib	1422.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	13	32	ib	1430.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	14	40	ib	1438.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	15	48	ib	1446.00	8	8	0	man	0	0
<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	16	56	ib	1454.00	8	8	0	man	0	0



概要

Celestial Sphere



North Celestial Pole

Declination does not change
 $\alpha = 5^{\circ} 23m$

→ path the Orion Nebula takes

Rises at 3:37 pm

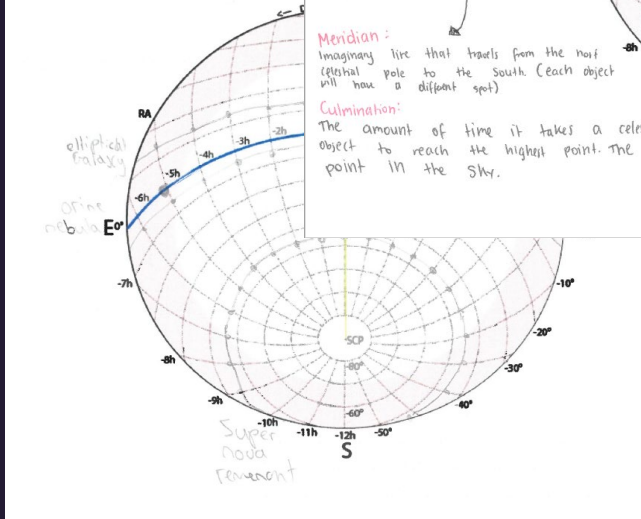
Sets at 10:37 pm

Orion Nebula

When an object crosses the meridian it is at its highest point in the sky

Meridian: Imaginary line that travels from the north celestial pole to the south (each object will have a different spot)

Culmination: The amount of time it takes a celestial object to reach the highest point in the sky.



Tracking objects: Right Ascension and Declination

Of the two coordinate systems to track objects Right Ascension and Declination uses time to locate objects. Below is a list of the 10 brightest radio objects from the Parkes Radio Telescope using RA/Dec. You can use this table, or move to a larger table with more objects here

Note: If you are in a class, choose an object that has not been selected by the majority. We will come together as a class to determine the most ideal object to view.

Name PKS Catalogue Number	Flux	Right Ascension (Time)			Declination (Degrees)			Culmination (Wait Time)			Time of Highest Point		
		h	m	s	d	m	s	h	m	s	h	m	s
Orion Nebula (PKS J0535-0523)	410000	05	35	18.40	-05	23	06.0	16	02	05	37		
RF 207 (PKS J1847-0202)	330000	18	47	18.10	-02	02	40.0	28	14				
Radio-Source (PKS J0859-4345)	260000	08	59	28.30	-43	45	44.0						
Supernova Remnant (PKS J0859-4731)	250000	08	59	07.30	-47	31	23.0	19	26	36			
Elliptical Galaxy (PKS J1230+1223)	220000	12	30	49.42	+12	23	28.0	22	57	10	32		
NA (PKS J1634-4735)	87000	16	34	36.40	-47	35	56.0	05		19	55		
NA (PKS J1346-6022)	79000	13	46	57.60	-60	22	58.0						
Pictor A (PKS J0519-4546)	66000	05	19	44.30	-45	46	40.0	16	06	4	19		
SNR M1ne 23 (PKS J1211-5809)	62000	12	11	56.00	-53	09	37.0						
Active Galactic Nuclei (PKS J1720-0058)	54000	17	20	30.40	-00	58	38.0	3	52	14	22		
NA (PKS J1639-4651)	53000	16	39	48.60	-46	51	01.0						

Preparation (Finding our time and place)

To track an object using time, we must use a fixed reference of time called the Local Apparent Sidereal Time. This time is calculated based on the Longitude of your position on Earth, and the First Point of Aries (Where the Sun crosses the Celestial Equator, historical name).

Tallaroek Observatory Longitude: 145.124958 E

D.L.R. Hosted Calculator

Question 1. Determine the Local Sidereal Time (LAST) using the link above, and entering the Longitude

Local Sidereal Time (LAST) = 11 hours 33 minutes 22 seconds

Current Time = 11 hours 25 minutes 24 seconds

Which object is visible during your observation window? Let's find out...

For your selected object we will determine the current Right Ascension and Declination to determine its place in the sky

11:33 : 22 sec
hr min

Responses cannot be edited

Celestial Object

You have selected an object, you have placed it in the sky, lets conduct some research whilst we're waiting for our Radio Map to be created using the radio telescope

*Required

What object did you choose? *

Write down the type and the name of the object you are observing

Supernova remnant

What is the RA and Declination of your object? *

RA: 08h 59m 07.3s DEC: -47d 31m 23.0s

What is the Culmination of your object? *

You can calculate this again using the Local Sidereal Time and the RA of the object. Remember, Culmination is the time it will take to reach the highest point.

19h 25m 32.3s

What time will it reach the highest point? *

30h 58m 12.3s

Research your chosen object *

You may want to include what the object is, why it is a strong radio source, what we already know about the object itself or the

Right Ascension

The Culmination of your object is the time it will take to reach the highest point in the sky, the telescope's meridian. Note: If RA < LAST add 24 hours to RA of object.

Remember, RA operates in a Sidereal Day (one rotation of the Earth) in 23 hrs 56 mins 4 secs... not 24 hours...

Question 2. Calculate the Culmination of your chosen object, show your working.

An approximate calculation is good enough!

Culmination = RA (Object) - LAST

$$24.19 (05.24 A) - 15.38 = 15.86$$

$$8:35:18.4 (\text{Orion nebula}) - 13:38:36$$

$$8:59:47.3 (\text{Supernova Remnant}) - 15:33:36 = 19.26 \text{ hr}$$

Add your calculated Culmination to the previous table, and share with the class

Is your object close to Culmination during your observation time?

Time of Highest Point = Current Time + Culmination (24 hour time)

$$\text{Orion } 10:36 + 16:02 = 27:38 \quad 27:38 - 24 = 3:37$$

$$\text{Supernova } 11:35 + 19:26 = 30:61 \quad 30:61 - 24 = 6:01 = 7:01$$

$$\text{Pictor A } 11:35 + 15:86 = 27:21$$

Declination

When will the object rise? Will it be high enough above the horizon to be observed?

There is no easy answer..... consider tracking the object in Stellarium with the Equatorial grid to investigate why

Declination	Notes
-80° > -90°	These objects are high enough all the time. But there are no interesting objects that close to the South Celestial Pole unfortunately...
-35° > -80°	These objects will disappear below the horizon at some time but they are always to the south of the telescope.
-34° > +15°	These objects are observable provided all of the observation is done either before they reach culmination or all is done after culmination because this point passes to the North of the telescope.
DEC > +15°	These objects are never high enough above the horizon for long enough to be useful targets from the Tallaroek Observatory. (View these from the Northern Hemisphere.)

As a class

In the table on the first page, add the classes calculations for Culmination and a Time of Highest Point of each object. It is recommended to observe the object that is the closest to your booked observation session.

Responses cannot be edited

Intro to Radio Astronomy

*Required

What is Radio Astronomy?

radio astronomy is a major branch of astronomy and reveals otherwise-hidden characteristics of everything in the universe. It studies celestial objects at radio frequencies.

What are some of the benefits to Radio Astronomy?

large and multiple field-of-views, rapid electronic steering, reliability, flexibility, cost and performance

How are these different to Radio Astronomy and Optical Astronomy?

Radio astronomy has some advantages over optical astronomy. Radio astronomy can be performed any time, day or night, since visible light from the sun does not interfere with radio emissions from other objects in the sky. Optical astronomy can only be done after sundown. Another advantage is that radio astronomy does not depend as much on atmospheric conditions. Radio waves are not blocked by dust or water vapor (clouds) in the atmosphere, so radio astronomy can be done in the rain. Optical astronomy requires a clear night. The only time radio astronomy cannot be performed is when there is lightning.

Why does a Radio Telescope need an Amplifier?

A receiver and amplifier to boost the very weak radio signal to a measurable level. These days the amplifiers are extremely sensitive and are normally cooled to very low temperatures to minimise interference due to the noise generated by the movement of the atoms in the meta

Answer the following True False Questions *

	True	False
Gamma Rays can reach Earth's Surface	<input type="checkbox"/>	<input checked="" type="checkbox"/>
All radio waves can reach the Earth's Surface	<input type="checkbox"/>	<input checked="" type="checkbox"/>
Satellites can be used to detect visible light	<input checked="" type="checkbox"/>	<input type="checkbox"/>

Sagittarius A East SNR G000.0+00.0

Type: supernova remnant (S)
RA/Dec (J2000.0): 17h45m41.04s/-29°00'46.8"
RA/Dec (on date): 17h47m00.46s/-29°01'14.2"
HA/Dec: 5h20m59.78s/-29°01'14.2"
Az./Alt.: +251°17'34.0"/+24°30'45.3"
Gal. long./lat.: -0°03'29.9"/-0°03'07.1"
Supergal. long./lat.: -174°09'22.3"/+42°14'41.6"
Ecl. long./lat. (J2000.0): +266°51'19.8"/-5°36'46.0"
Ecl. long./lat. (on date): +267°08'46.9"/-5°36'55.9"
Ecliptic obliquity (on date): +23°26'14.1"
Mean Sidereal Time: 23h07m00.2s
Apparent Sidereal Time: 23h07m59.2s
Rise: 1h37m
Transit: 9h20m
Set: 17h04m
Parallactic Angle: +121°14'56.0"
IAU Constellation: Sgr
Size: +0°03'30.00" x +0°02'30.00"
Morphological description: remnant shows a shell radio structure.

Canopus

Sun

Venus

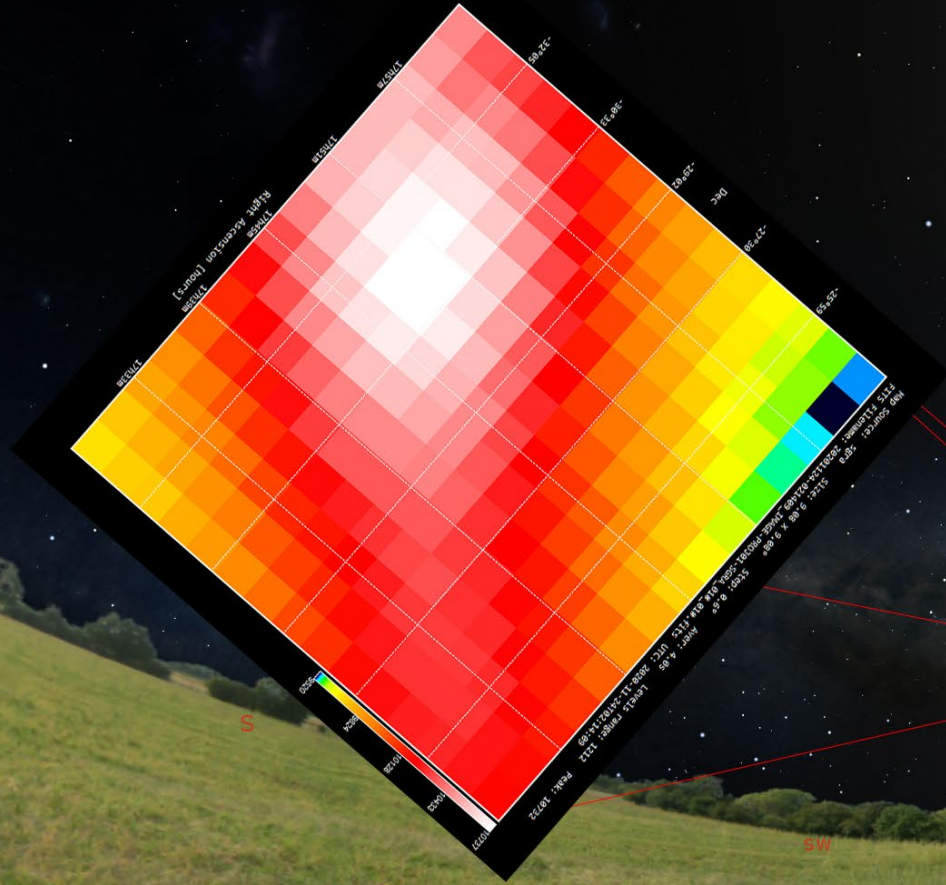
Jupiter

Mercury

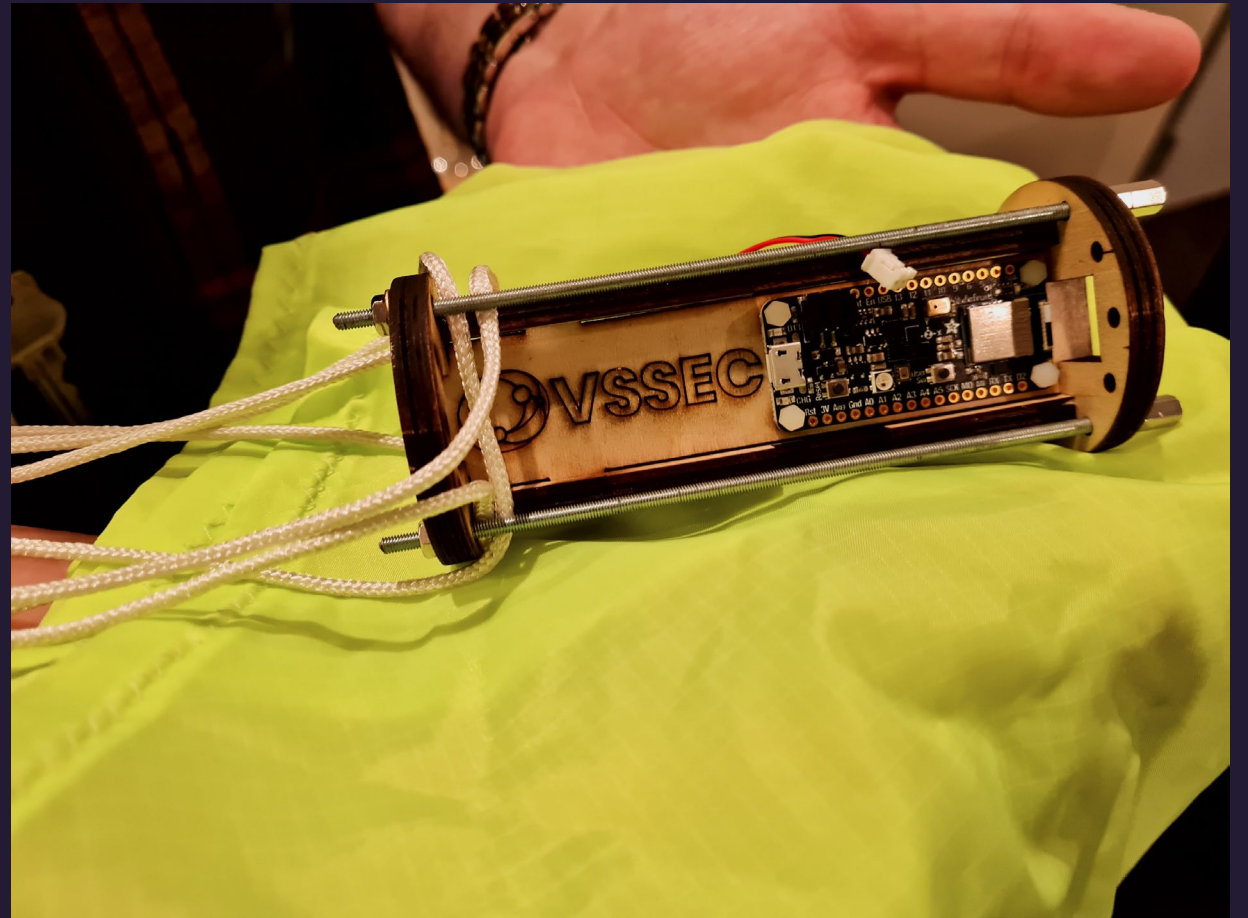
Saturn

Altair

Antares



学校でのVSSECの缶サット (CanSat)





缶サットを用いた教育の流れ

教室の雰囲気
づくり

衛星調査・
技術の進歩

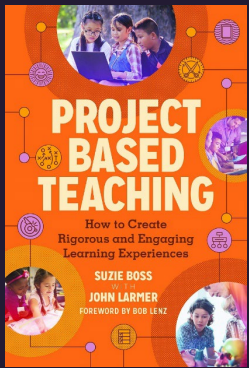
パラシュート
の設計と試験

シャーシの
設計と試験

プログラミング
と電子機器の
開発と試験

データ解析と
プレゼンテー
ション

打ち上げ!



European Space Agency
Agence spatiale européenne

学習管理システム

VSSEC Moodle

CanSAT (MASTER Course)

Dashboard / Courses / Miscellaneous / CanSAT

Turn editing on

The CanSat project will give students practical experience in electronics, computer programming, remote sensing and data analysis.








This project extends skills in the [Engineering Design Process](#) and Scientific Method, using Project Based Learning. Students and teachers are challenged to integrate Science, Technology, Engineering, Arts/Design, and Mathematics (STEAM).

Students will gain the confidence to undertake project based work and will be involved with other students through collaborative means. Students are supported, encouraged to take risks, and celebrate their mistakes and triumphs.

The project is a one semester course undertaken as part of a student's normal class time. However, you are encouraged to modify the program to suit your needs and that of your classroom.

If you are new to this program, start by clicking "**Introduction**" below.

Announcements

 Introduction	 Unit 0: Building Classroom Culture	 Unit 1: Your Mission	 Unit 2: Building the CanSAT
 Unit 3: CanSAT Coding and Electronics	 CanSAT Launch!	 Unit 4: Data Analysis and Presentation	